PAPER • OPEN ACCESS

Radiative neutrino mass generation from WIMP dark matter

To cite this article: Roberto A Lineros 2016 J. Phys.: Conf. Ser. 718 042036

View the <u>article online</u> for updates and enhancements.

Related content

- Mass generation and the problem of seagull divergences
 C. T. Figueiredo and A. C. Aguilar
- Radiative neutrino mass model with a TeV scale gauge symmetry
 Daijiro Suematsu
- Radiative neutrino mass generation and dark energy K Bamba, C Q Geng and S H Ho

Radiative neutrino mass generation from WIMP dark matter

Roberto A Lineros

Instituto de Física Corpuscular – CSIC/U. Valencia, Parc Científic, calle Catedrático José Beltrán 2, E-46980 Paterna, Spain

E-mail: rlineros@ific.uv.es

Abstract. The minimal seesaw extension of the Standard Model requires two electroweak singlet fermions in order to accommodate the neutrino oscillation parameters at tree level. Here we consider a next to minimal extension where light neutrino masses are generated radiatively by two electroweak fermions: one singlet and one triplet under SU(2). These should be odd under a parity symmetry and their mixing gives rise to a stable weakly interactive massive particle dark matter candidate. For mass in the GeV-TeV range, it reproduces the correct relic density, and provides an observable signal in nuclear recoil direct detection experiments. The fermion triplet component of the dark matter has gauge interactions, making it also detectable at present and near future collider experiments.

1. Introduction

It is well established that most of the content of the Universe is in the form of Dark Matter (DM). The latest observations of the Cosmic Microwave Background anisotropies [1] implies that the DM relic abundance in the Λ CDM model is

$$\Omega_{\rm DM}h^2 = 0.1198 \pm 0.0015\,, (1)$$

where $h = 0.678 \pm 0.009$ is the scale factor for Hubble expansion rate [1]. The value of the relic abundance can be explained in terms of Weakly Interactive Massive Particles (WIMPs). This class of particles are not part of the Standard Model (SM) indicating the necessity to go beyond it. The presence of DM motivates in part searches of new particles in particle accelerator like the Large Hadron Collider (LHC).

In general words, WIMP-DM particles can easily reproduce the observed value of the relic abundance thanks to the *freeze-out mechanism*. This mechanism comes from the combined effect of the expansion of the Universe and the interaction rate among DM and the primordial plasma during the early Universe.

For WIMP particles in GeV-TeV range, the relic abundance is

$$\Omega_{\text{WIMP}} h^2 \simeq 0.1 \frac{3 \times 10^{-26} \,\text{cm}^3/\text{s}}{\langle \sigma v \rangle},$$
(2)

where $\langle \sigma v \rangle$ is the DM thermally averaged cross section. The correct value of $\Omega_{\rm DM} h^2$ is then obtained when $\langle \sigma v \rangle$ is of the order of weak interaction cross sections. In minimal scenarios, DM

Content from this work may be used under the terms of the Creative Commons Attribution 3.0 licence. Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI.

	Standard Model			Fermions		Scalars	
	L	e	ϕ	Σ	N	η	Ω
$SU(2)_L$	2	1	2	3	1	2	3
Y	-1	-2	1	0	0	1	0
Z_2	+	+	+	—	_	-	+

Table 1. Matter content.

models are constructed on top of the SM providing also new mediators to act as portal between the dark and visible sectors. However, it seems reasonable that DM is part of dark sector that includes new particles and interactions like in Supersymmetric, extra dimensions [2], or dark gauge group [3] models.

On the other hand, neutrinos also give an indication that the SM is not complete. In the SM, neutrinos are massless but the observation of neutrino oscillations indicate the contrary, i.e. they are massive and with the largest mass squared differences of $\sim 10^{-3}\,\mathrm{eV}^2$ [4, 5]. The most known mechanism of neutrino mass generation are the so called see-saw models (see for instance [6] and references within). Where the smallness of neutrino masses is generated at tree level by extremely heavy states. However, it is also possible to generate the neutrino masses via loop processes. One realization are the scotogenic models [7, 8]. In one version, the SM is enriched with an extra $SU(2)_L$ doublet scalar and a majorana fermion singlet. In the second one, it is instead added an extra $SU(2)_L$ doublet scalar but a $SU(2)_L$ fermion triplet. In this type of models, an ad-hoc symmetry is imposed and therefore mass generation at tree level is forbidden. This scheme is very tantalizing due to that the ad-hod symmetry allows to have a DM candidate.

Our model [9] is based on a mixed realization of the scotogenic models. We show that in this scheme the phenomenology is richer and allow to have a WIMP-DM candidate with a sizeable interaction with quarks and a mass ranging from GeV to TeV.

2. The model

As previously stated, the model is a mix realization of two scotogenic models. Besides the particle content of the SM, we have included a majorana fermion singlet N, a $SU(2)_L$ fermion triplet Σ , a $SU(2)_L$ scalar triplet Ω ; the latter three are without hypercharge, and finally a $SU(2)_L$ scalar doublet η . The new fields, except Ω , are charged with a Z_2 symmetry that ensures the DM stability. Due to Ω is not protected by the symmetry, it will participate in the spontaneous symmetry breaking of the $SU(2)_L \times U(1)_Y$ SM group. All the matter content of the model is summarized in Tab. 1.

Using the symmetries described in the matter content, we write the lagrangian of the model

$$\mathcal{L} \supset -Y_{\alpha\beta} \overline{L}_{\alpha} e_{\beta} \phi - Y_{\Sigma_{\alpha}} \overline{L}_{\alpha} C \Sigma^{\dagger} \tilde{\eta} - \frac{1}{4} M_{\Sigma} \text{Tr} \left[\overline{\Sigma}^{c} \Sigma \right] +$$

$$-Y_{\Omega} \text{Tr} \left[\overline{\Sigma} \Omega \right] N - Y_{N_{\alpha}} \overline{L}_{\alpha} \tilde{\eta} N - \frac{1}{2} M_{N} \overline{N}^{c} N + h.c. - V_{\text{scal}},$$
(3)

where Y_X stand for yukawa couplings, M_X are mass terms, and V_{scal} is the scalar potential. The details about the scalar potential are in Ref. [9].

The neutrino masses arises at one loop mainly due to the terms $L\eta N$ and $L\Sigma \eta$. We obtain

$$M_{\alpha\beta}^{\nu} = \sum_{\sigma=1,2} \frac{h_{\alpha\sigma} h_{\beta\sigma}}{8\pi^2} \frac{\lambda_5 v_h^2}{m_0^2} M_k , \qquad (4)$$

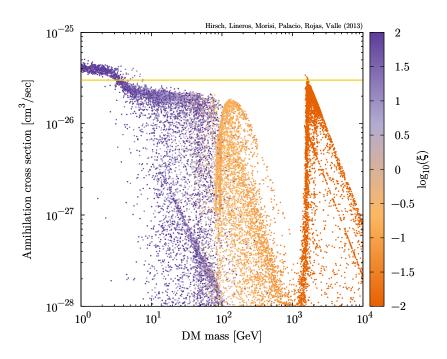


Figure 1. Annihilation cross section versus DM mass. More details in Ref. [9]

where $h_{\alpha\beta}$ is a 3 × 2 matrix that encodes the values of the yukawa terms Y_{Σ} and Y_N . v_h is the vacuum expectation value of the doublet ϕ , and M_k the masses of the mass eigenstates of the neutral fermions arise from N and Σ^0 (neutral component of the triplet Σ).

The model contains two possible DM candidates. One is the neutral component of the scalar doublet η which has a similar phenomenology to the inert higgs model; and the second is the fermion which is a combination among N and Σ^0 . In our analysis, we focus on the fermion DM candidate. The two neutral fermions mass eigenstate are combination of N and Σ^0 and we will denote as $\chi_{1,2}^0$. This class of mixed state, between singlet and triplet, was not possible in the traditional versions of the scotogenic model. This is possible thanks to Ω that allows the term $\Omega \Sigma N$ in the lagrangian. This term also generates a mass splitting between the neutral and charged component of the triplet Σ which later it will reveal to be essential for the DM phenomenology.

At this point, it is important to describe the mass spectrum of the fermion states:

$$m_{\chi^{\pm}} = M_{\Sigma}, \qquad (5)$$

$$m_{\chi_1^0} = \frac{1}{2} \left(M_{\Sigma} + M_N - \sqrt{(M_{\Sigma} - M_N)^2 + 4(2Y_{\Omega}v_{\Omega})^2} \right),$$
 (6)

$$m_{\chi_2^0} = \frac{1}{2} \left(M_{\Sigma} + M_N + \sqrt{(M_{\Sigma} - M_N)^2 + 4(2Y_{\Omega}v_{\Omega})^2} \right),$$
 (7)

$$m_{\chi_1^0} = \frac{1}{2} \left(M_{\Sigma} + M_N - \sqrt{(M_{\Sigma} - M_N)^2 + 4(2Y_{\Omega}v_{\Omega})^2} \right),$$

$$m_{\chi_2^0} = \frac{1}{2} \left(M_{\Sigma} + M_N + \sqrt{(M_{\Sigma} - M_N)^2 + 4(2Y_{\Omega}v_{\Omega})^2} \right),$$

$$\tan(2\alpha) = \frac{4Y_{\Omega}v_{\Omega}}{M_{\Sigma} - M_N},$$
(8)

where v_{Ω} is the vacuum expectation value of the neutral component of Ω and α the mixing angle between N and Σ^0 . Here it is easy to appreciate the role of v_{Ω} breaking the degeneracy between χ^{\pm} and the χ_i^0 with larger component of Σ^0 .

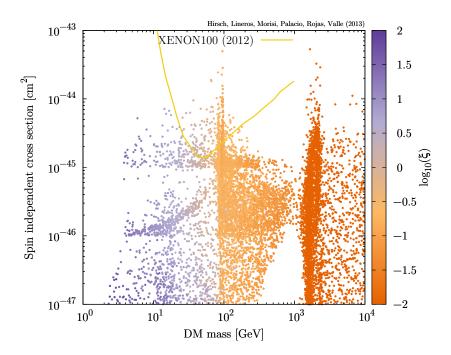


Figure 2. Spin-independent cross section versus DM mass. More details in Ref. [9]

We use the codes LanHEP [10, 11, 12] and MicrOmegas [13] to scan on the parameter space (See details in Ref. [9]) including constraints from neutrino masses, perturvativity limits of the adimensional parameters, and LEP [14] and CMS [15] searches on new charged particles. Also, we require that $m_{\chi_1^0} < m_{\chi^\pm}, m_{\eta}$. We obtained the parameter space compatible with current relic abundance value (Eq. 1). The allowed region is compatible with DM masses in the range of 10 to 10^4 GeV.

In Fig. 1, we present the result of the scan in terms of the annihilation cross section (relevant for indirect detection searches) versus the DM mass. The color code corresponds to the value of

$$\xi = \frac{M_{\Sigma} - m_{\rm DM}}{m_{\rm DM}} \,, \tag{9}$$

which is related to the amount of Σ^0 present in DM candidate. This quantity reveals three regimes. The low DM mass corresponds to a DM which is mainly singlet N. This is the consequence of the LEP constraints on charge particles. At the TeV masses, the DM candidate is mostly triplet. This is produced due to annihilation channels to gauge bosons and coannihilation with χ^{\pm} dominates. In the range between 100 GeV to 1 TeV, the DM is more a mixed state N- Σ^0 . This regime is very interesting because annihilations into quarks become dominating. This occurs because $\Omega \Sigma N$ allows a higgs portal.

Same three regimes can be observed in Fig. 2 where we show the spin-independent cross section versus DM mass. Let us highlight that in traditional scotogenic models the expected spin-independent cross section is largely suppressed because DM can only interact with quarks via loops. In addition, the direct detection signal is of the order of the reach of current experiments. In Fig. 2, we also show latest exclusion limits of XENON100 [16] as reference.

3. Conclusions

The nature behind DM is still a mystery. WIMPs are the most attractive candidates due to many search strategies. We have proposed a new scotogenic model that provide a neutrino mass mechanism at one loop and a fermion WIMP-DM candidate. The fermion DM in this model provides features that are not present in the traditional scotogenic models. One feature is the non zero spin independent cross section that is at the reach of current and future direct detection experiment. Also, the DM mass can go from 10 GeV to 10 TeV which is broader with respect to previous models.

Acknowledgments

We would like to thank you the organizers of the TAUP 2015 for an interesting conference. This work was supported by the Spanish MINECO under grants FPA2014-58183-P, and MULTIDARK CSD2009-00064 (Consolider-Ingenio 2010 Programme); by Generalitat Valenciana grant PROMETEOII/2014/084, and Centro de Excelencia Severo Ochoa SEV-2014-0398. R. L. is supported by a Juan de la Cierva contract JCI-2012-12901 (MINECO).

References

- [1] Planck Collaboration 2015 Preprint arXiv:1502.01589
- [2] Lineros R A and Pereira dos Santos F A 2014 J. Cosmology Astropart. Phys. 10 059
- [3] França U, Lineros R A, Palacio J and Pastor S 2013 Phys. Rev. D 87 123521
- [4] Schwetz T, Tortola M and Valle J 2011 New J. Phys. 13 109401
- [5] Forero D, Tortola M and Valle J 2012 Phys. Rev. D86 073012
- [6] Lattanzi M, Lineros R A and Taoso M 2014 New Journal of Physics 16 125012
- [7] Ma E 2006 Phys.Rev. **D73** 077301
- [8] Ma E and Suematsu D 2009 Mod. Phys. Lett. A24 583-589
- [9] Hirsch M, Lineros R A, Morisi S, Palacio J, Rojas N and Valle J 2013 JHEP 10 149
- [10] Semenov A V 1996 Preprint arXiv:hep-ph/9608488
- [11] Semenov A V 2009 Computer Physics Communications 180 431
- [12]Semenov A 2010 Preprint arXiv:1005.1909
- [13] Belanger G, Boudjema F, Pukhov A and Semenov A 2013 Preprint arXiv:1305.0237
- $[14]\ \, \mathrm{L3}$ Collaboration 2001 Physics Letters B $\mathbf{517}\ 75\text{--}85$
- [15] CMS Collaboration 2012 Physics Letters B 718 348
- [16] Aprile E et al. 2012 Physical Review Letters 109 181301